

Einstein Metrics,
Four-Manifolds, &
Gravitational Instantons

Claude LeBrun
Stony Brook University

Canonical Metrics in Differential Geometry,
University of Wisconsin, Madison,
April 12, 2025.

Most recent results joint with

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Olivier Biquard
Sorbonne Université

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and

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(2024) 13295-13311.

Will also briefly mention

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work in progress

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Tristan Ozuch

MIT

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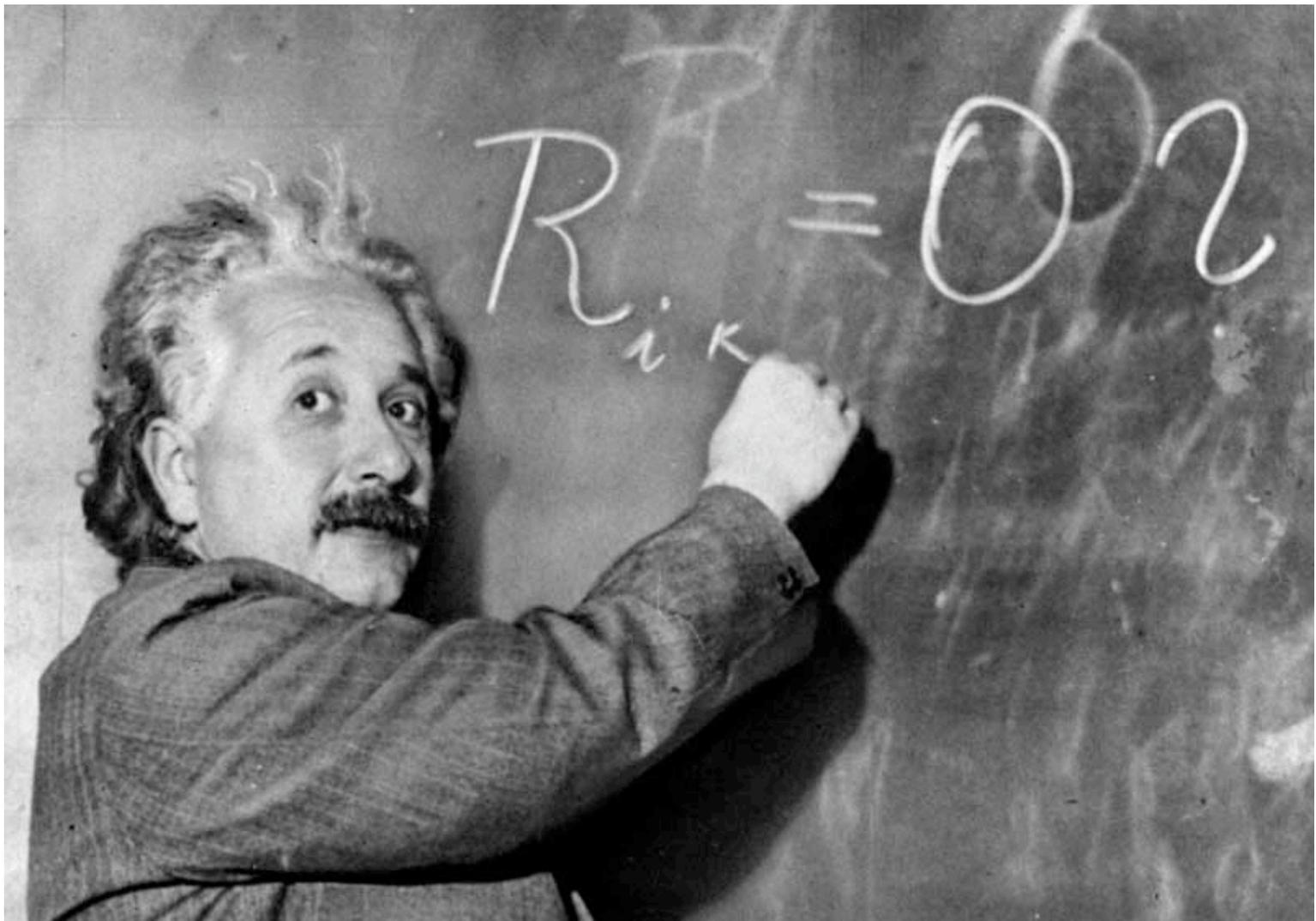
“... the greatest blunder of my life!”

— A. Einstein, to G. Gamow

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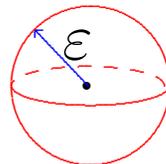
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$$\frac{\text{vol}_g(B_\varepsilon(p))}{c_n \varepsilon^n} = 1 - s \frac{\varepsilon^2}{6(n+2)} + O(\varepsilon^4)$$



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In high dimensions, the Einstein condition allows for a surprising degree of flexibility!

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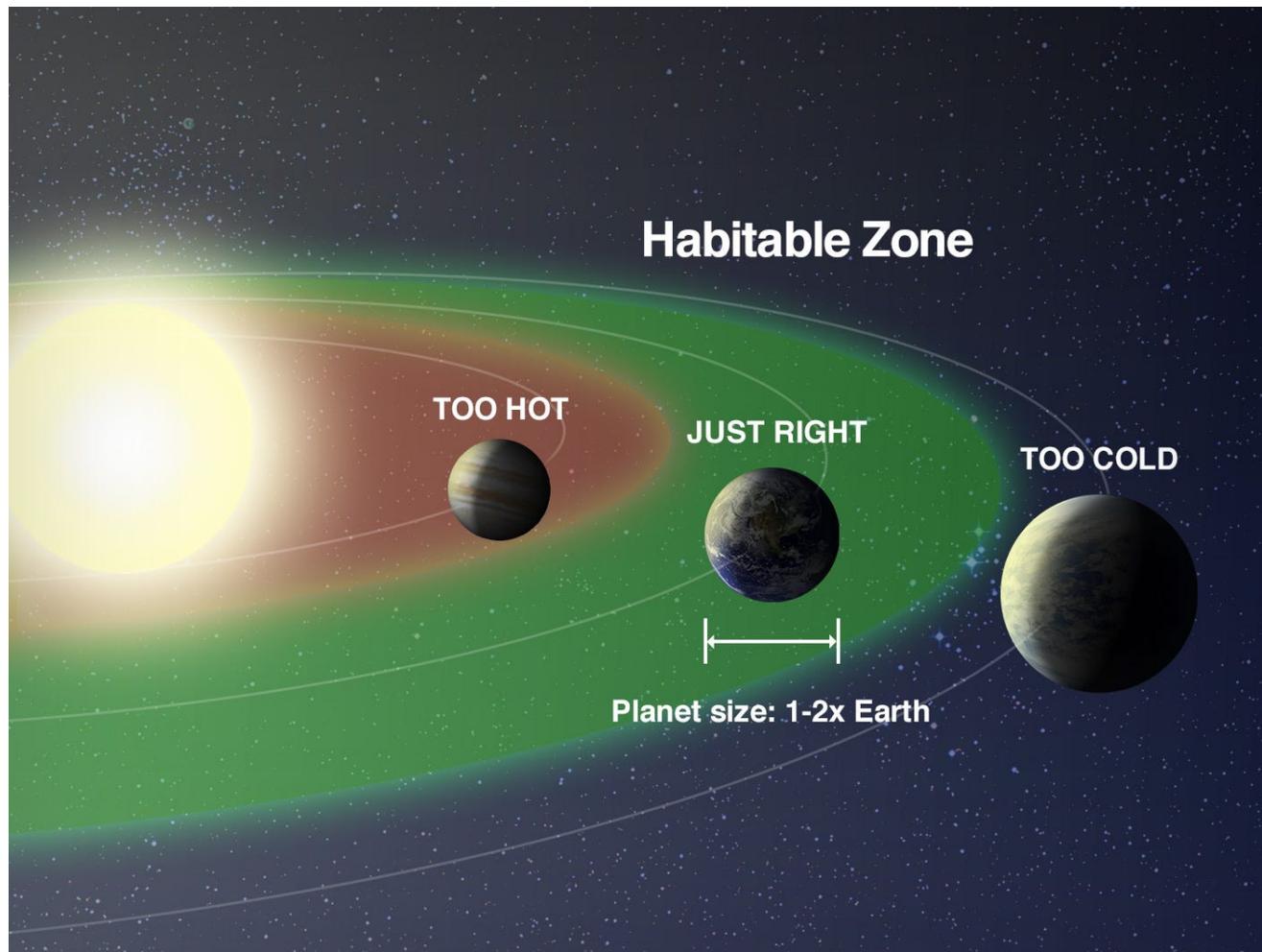
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Λ^+ self-dual 2-forms.

Λ^- anti-self-dual 2-forms.

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Which 4-manifolds admit Einstein metrics?

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$$d\omega = 0, \quad \lrcorner\omega : TM \xrightarrow{\cong} T^*M.$$

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$$\omega = dx \wedge dy + dz \wedge dt$$

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Some Suggestive Questions. *If (M^4, ω) is a symplectic 4-manifold, when does M^4 admit an Einstein metric g (a priori unrelated to ω)? What if we also require $\lambda > 0$?*

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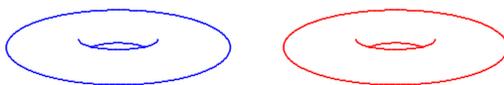
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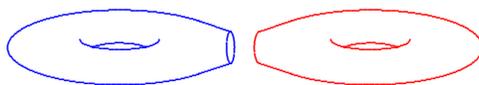
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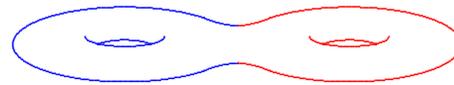
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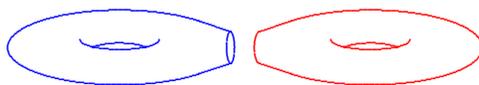
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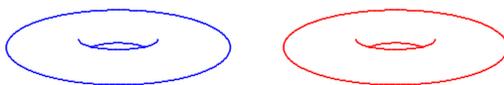
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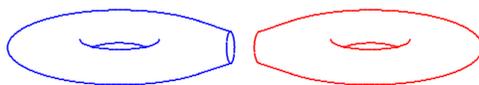
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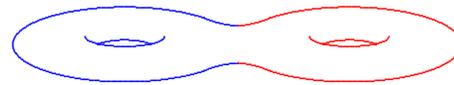
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Diffeotypes: exactly the Del Pezzo surfaces.

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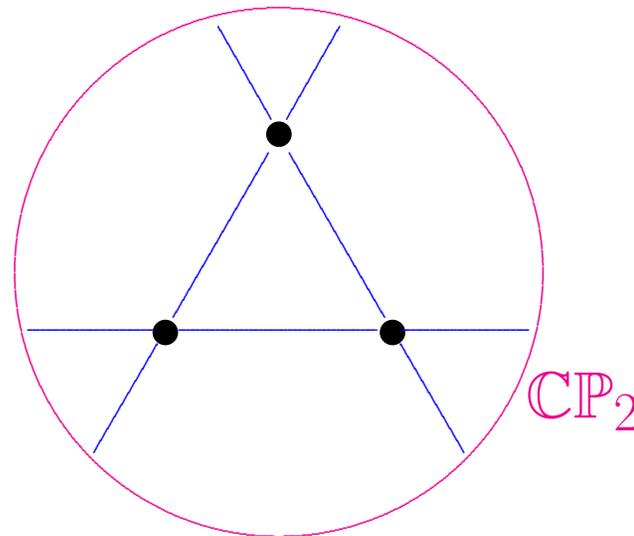
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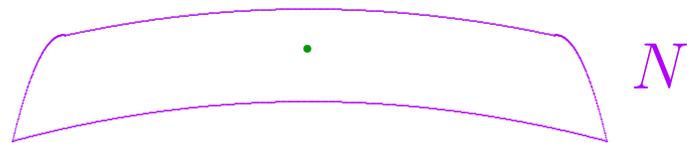
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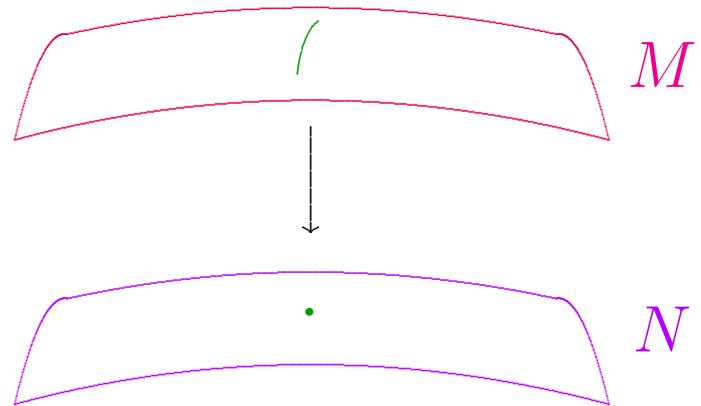
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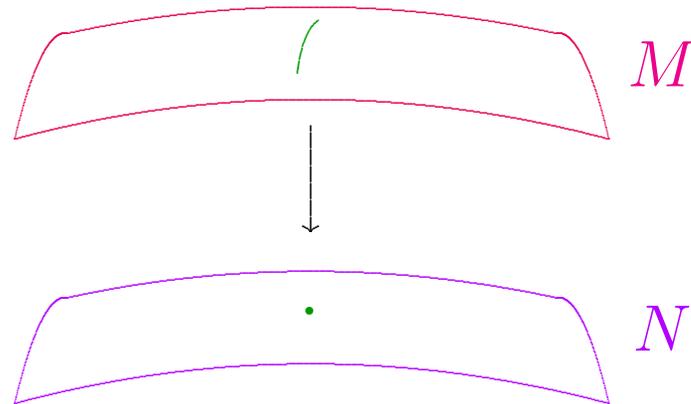
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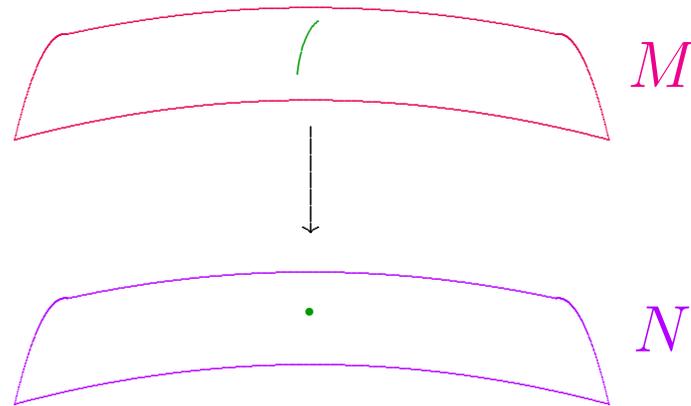


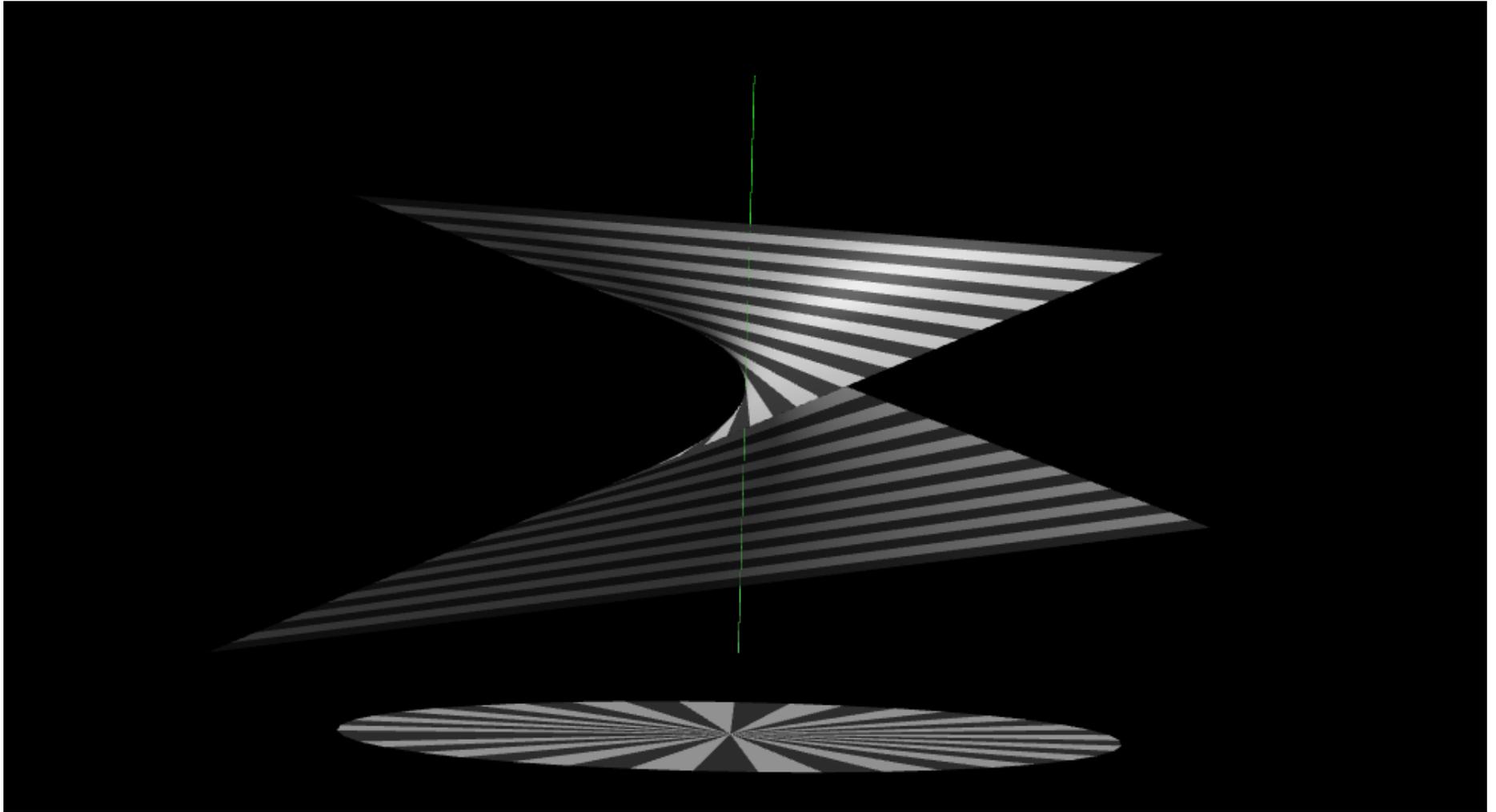
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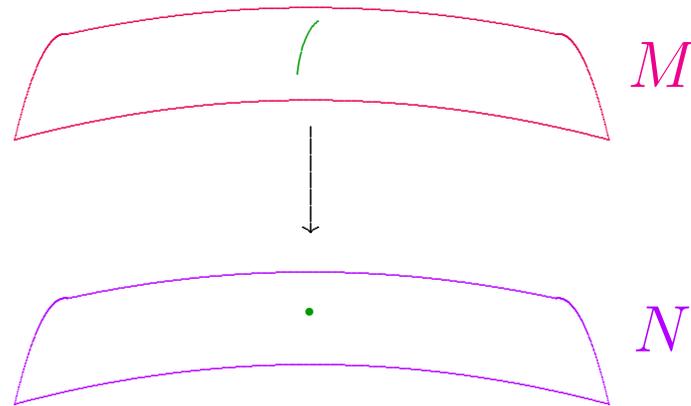


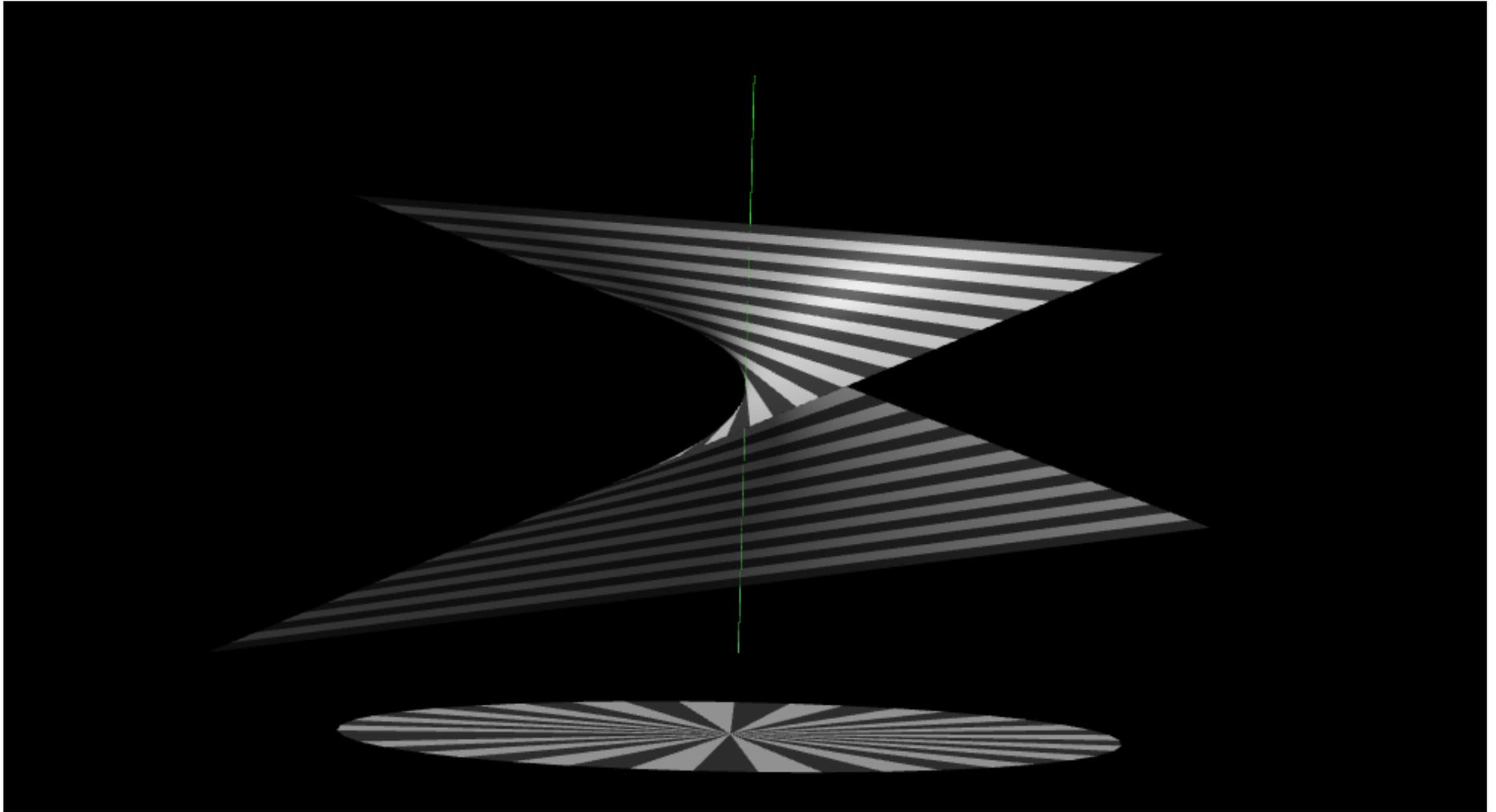
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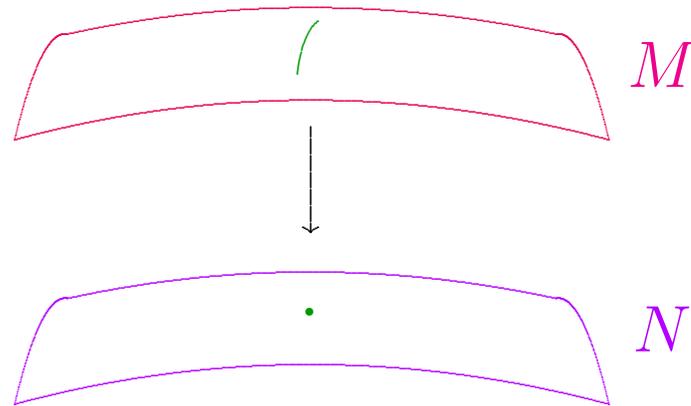


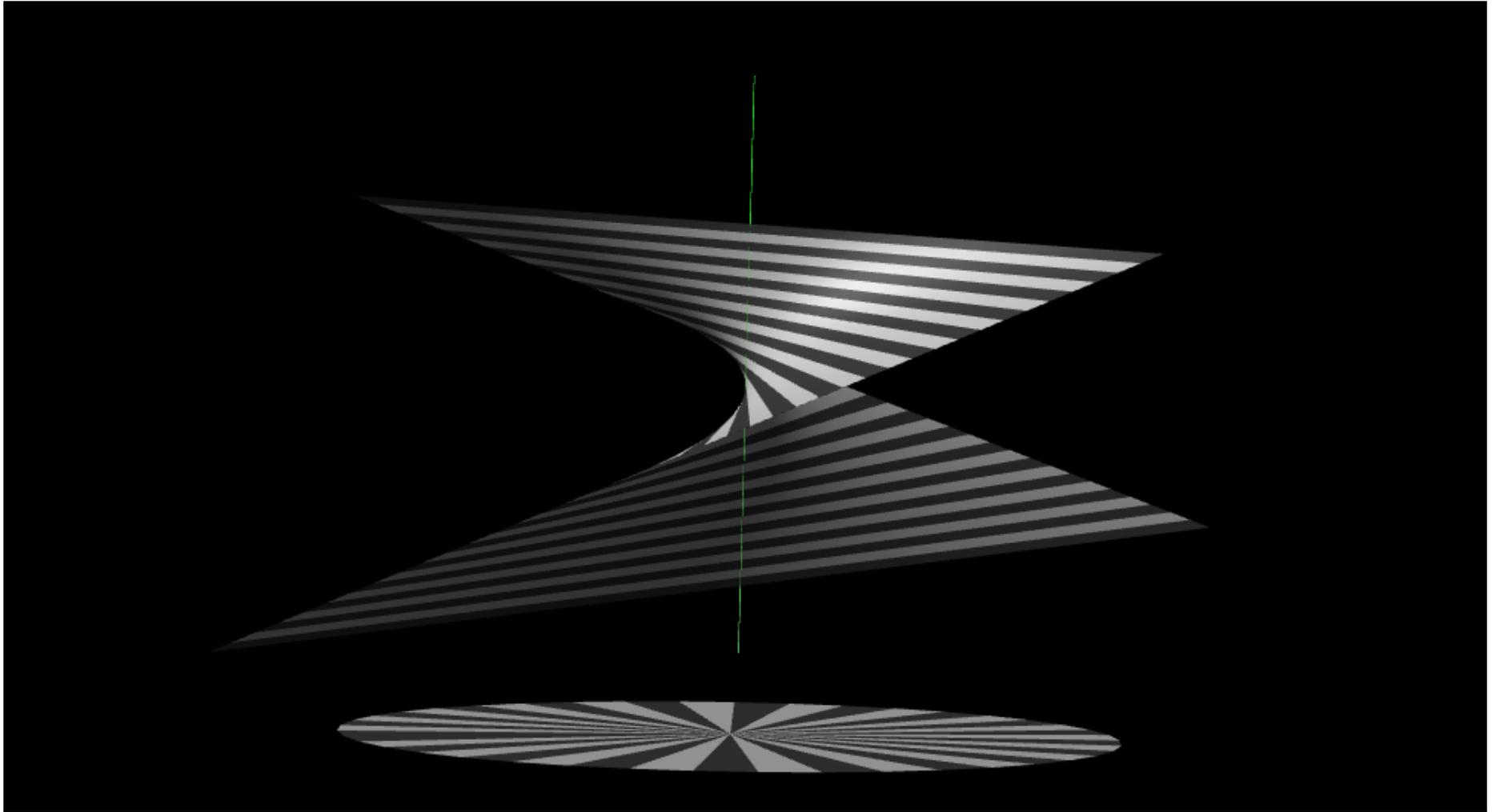
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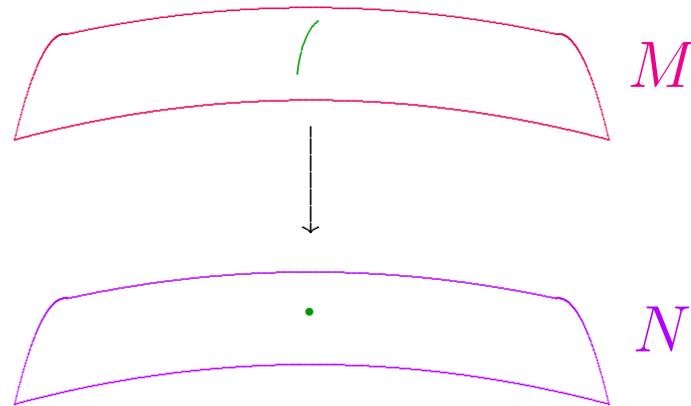


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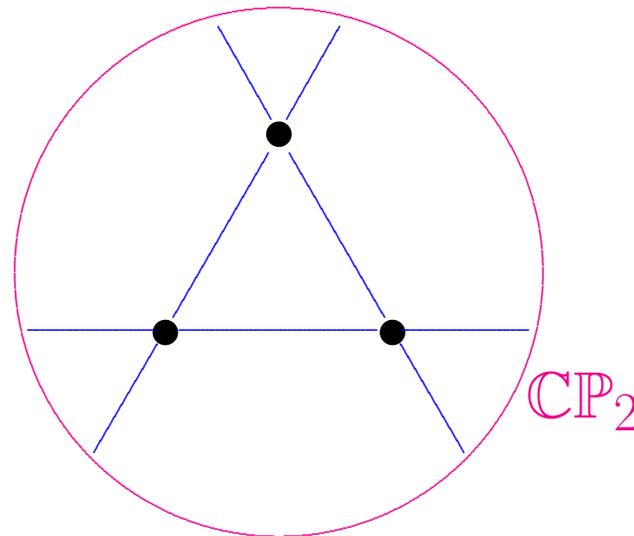


Del Pezzo surfaces:

(M^4, J) for which c_1 is a Kähler class $[\omega]$.

Shorthand: “ $c_1 > 0$.”

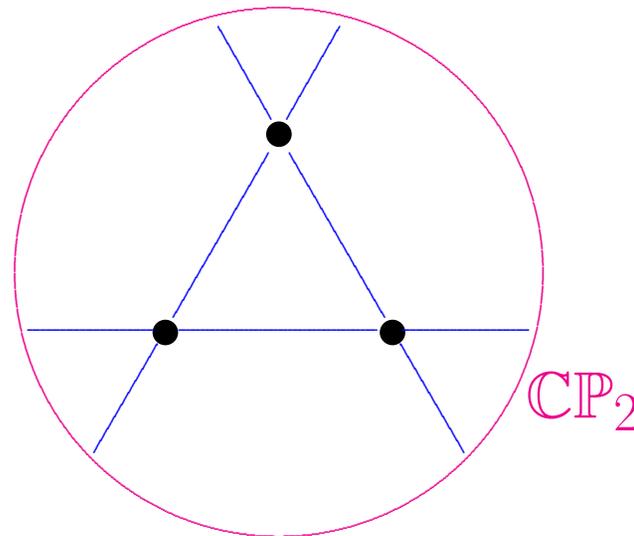
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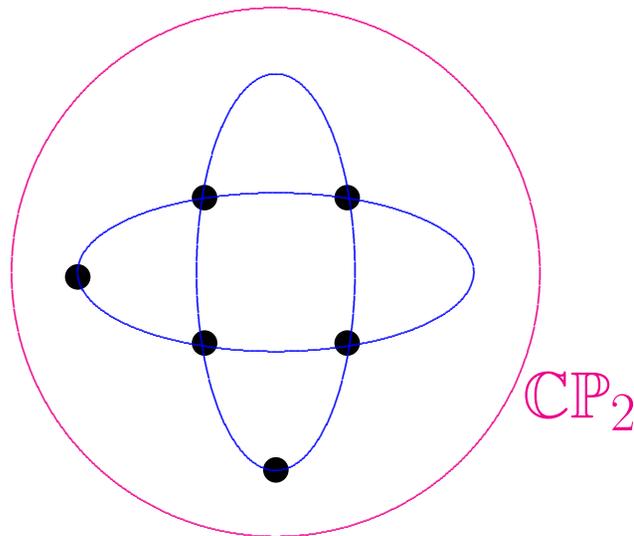


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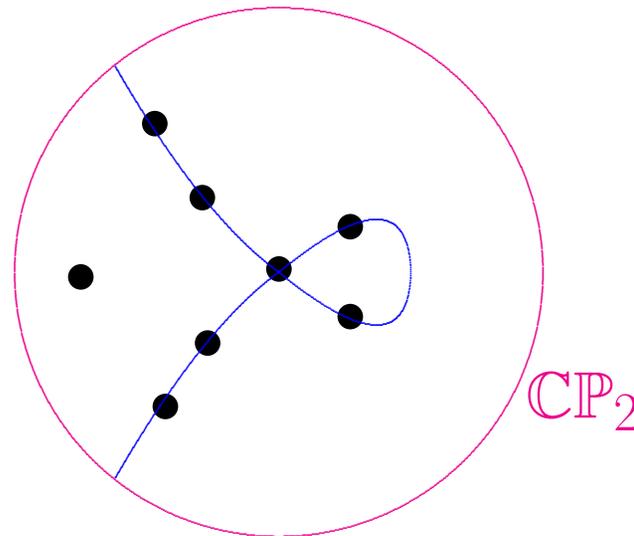


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$$g = g(J\cdot, J\cdot)$$

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Corollary. *Every simply-connected compact oriented Einstein (M^4, h) with $\det(W_+) > 0$ is diffeomorphic to a del Pezzo surface.*

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Corollary. *Every simply-connected compact oriented Einstein (M^4, h) with $\det(W_+) > 0$ is diffeomorphic to a del Pezzo surface. Conversely, every del Pezzo M^4 carries Einstein h with $\det(W_+) > 0$, and these sweep out exactly one connected component of moduli space $\mathcal{E}(M)$.*

In joint work with [Tristan Ozuch](#), currently extending these results to non-collapsed Gromov-Hausdorff limits of compact $\lambda > 0$ Einstein 4-manifolds.

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This illustrates how [gravitational instantons](#) play a crucial role, even when studying compact case.

Gravitational Instantons?

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Definition. *A gravitational instanton is a complete,*

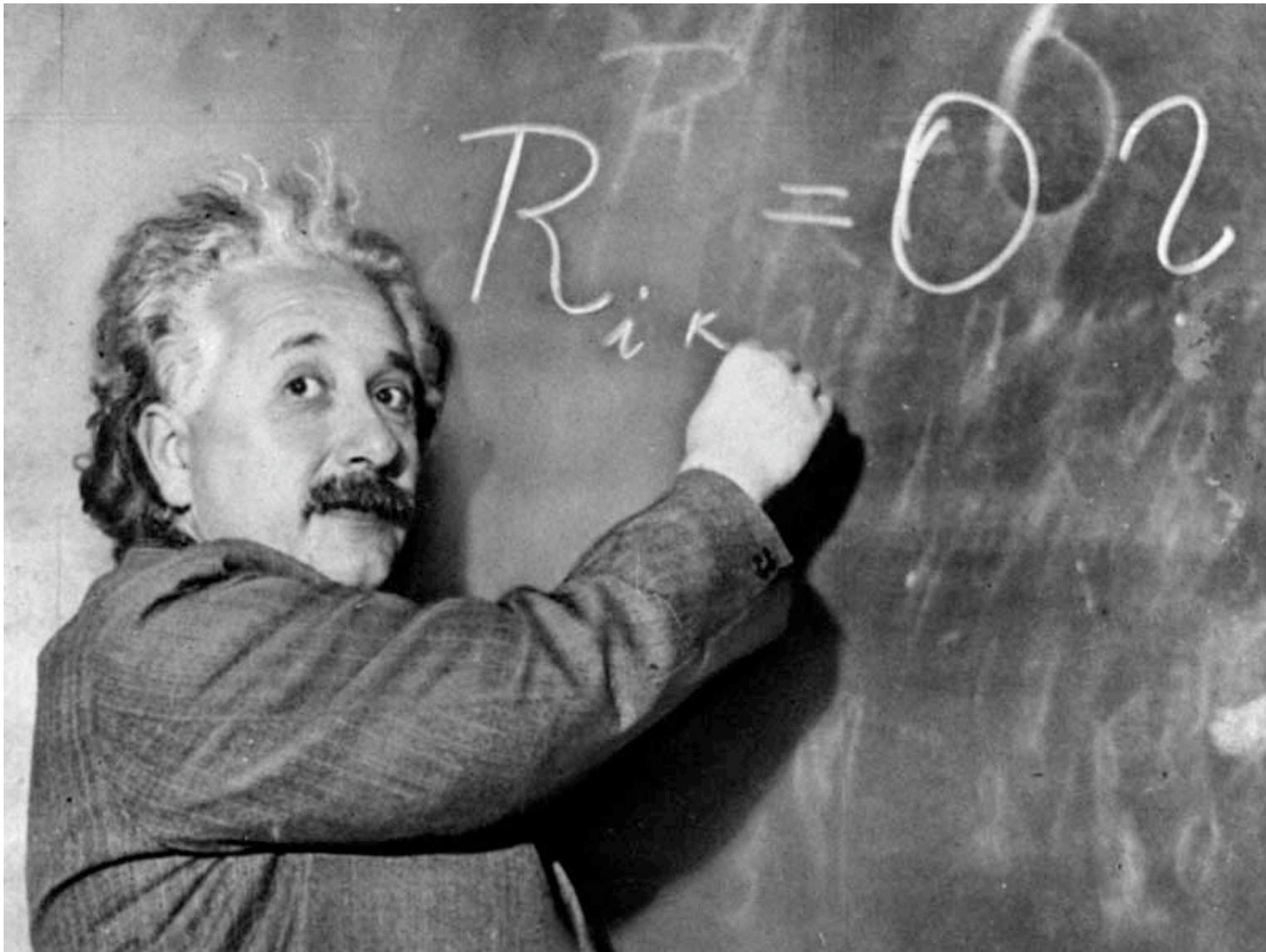
Definition. *A gravitational instanton is a complete, non-compact,*

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Key examples:

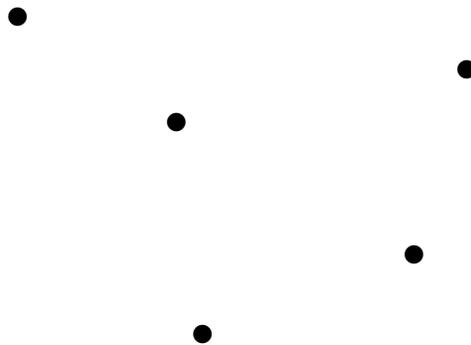
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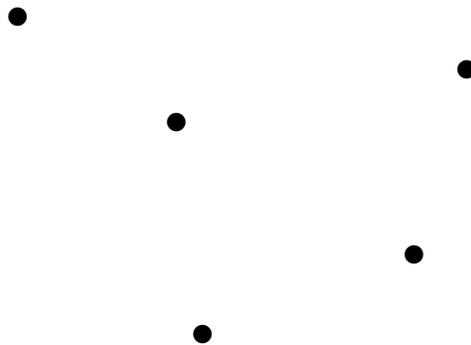
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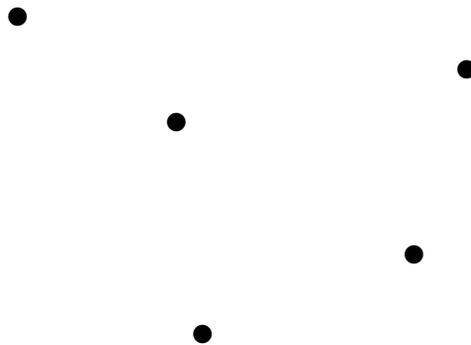
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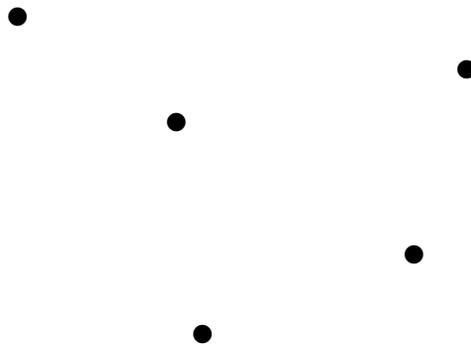


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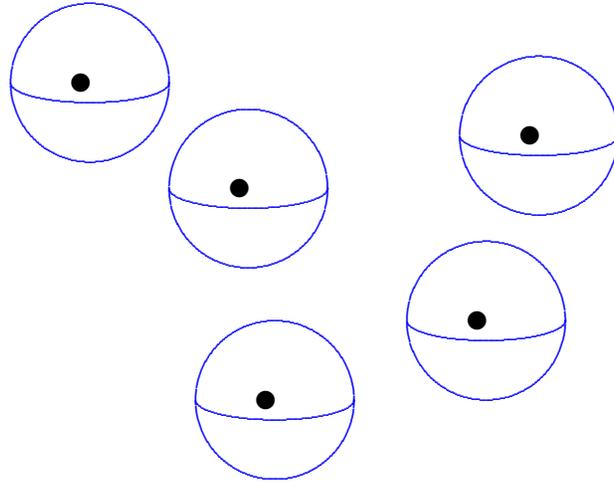
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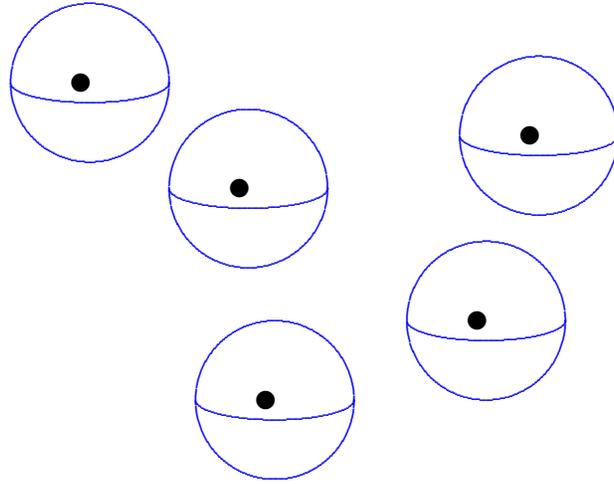
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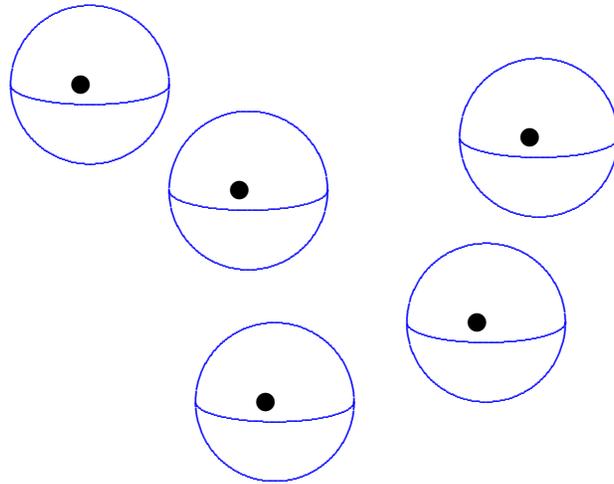
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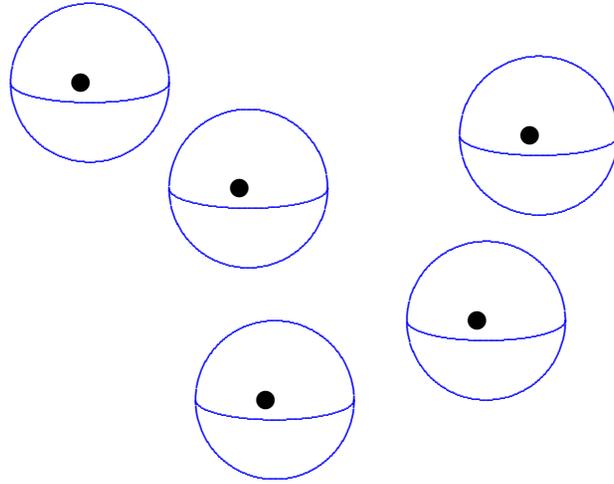
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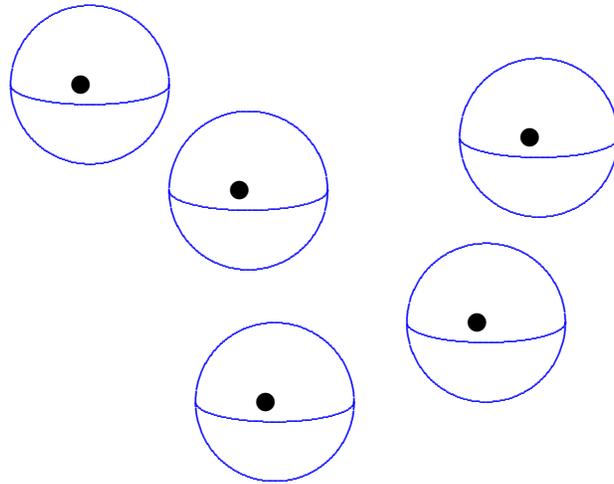
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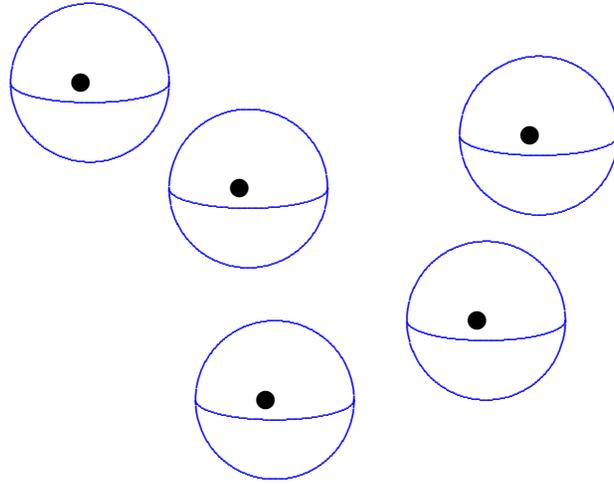
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This last property distinguishes the ALF spaces from other classes of gravitational instantons:

ALG, ALH, ALG*, ALH*, ...

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Non-Kähler, but **conformally** Kähler!

Hawking also explored non-hyper-Kähler examples. . .

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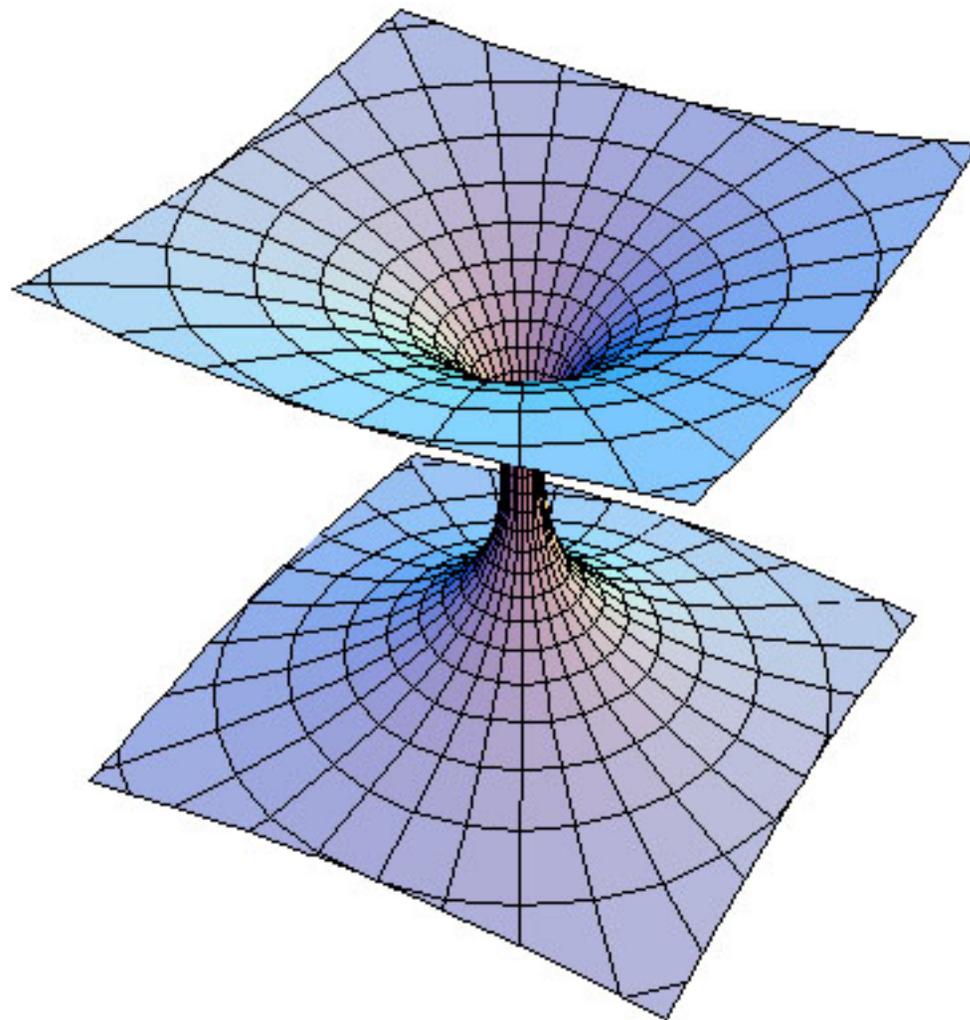
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$$\mathbb{R} \times S^2 \subset \mathbb{R}^2 \times S^2$$

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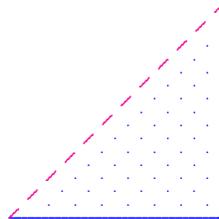
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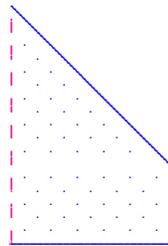
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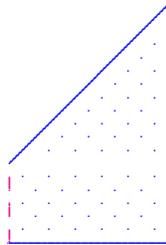
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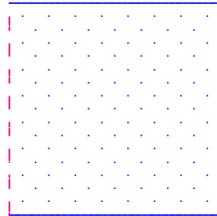
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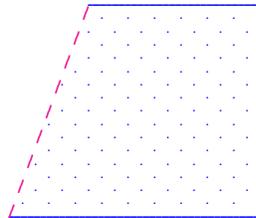
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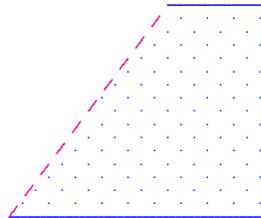
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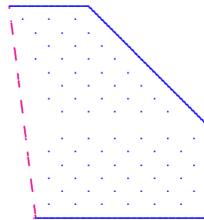
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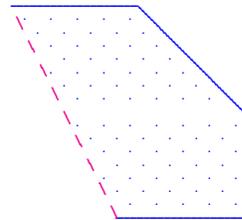
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Set $h = \alpha^{2/3}g$, where α top eigenvalue of W_{+g} , and choose top eigenform $\omega \in \Lambda^+$ with $|\omega|_h \equiv 1$. Then

$$0 \geq |\nabla\omega|^2 + 3\langle\omega, (d + d^*)^2\omega\rangle$$

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This optimal result combines **Theorem A** with a result of Mingyang Li, [arXiv:2310.13197](https://arxiv.org/abs/2310.13197).

Thanks for the invitation!

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